

Plasma Reactions of O_n High Energy Species – O , $O_2(v)$, $O_2(a^1\Delta_g)$, and O_3

O_2 plasma kinetics workshop
Reykjavik
Sep 2016



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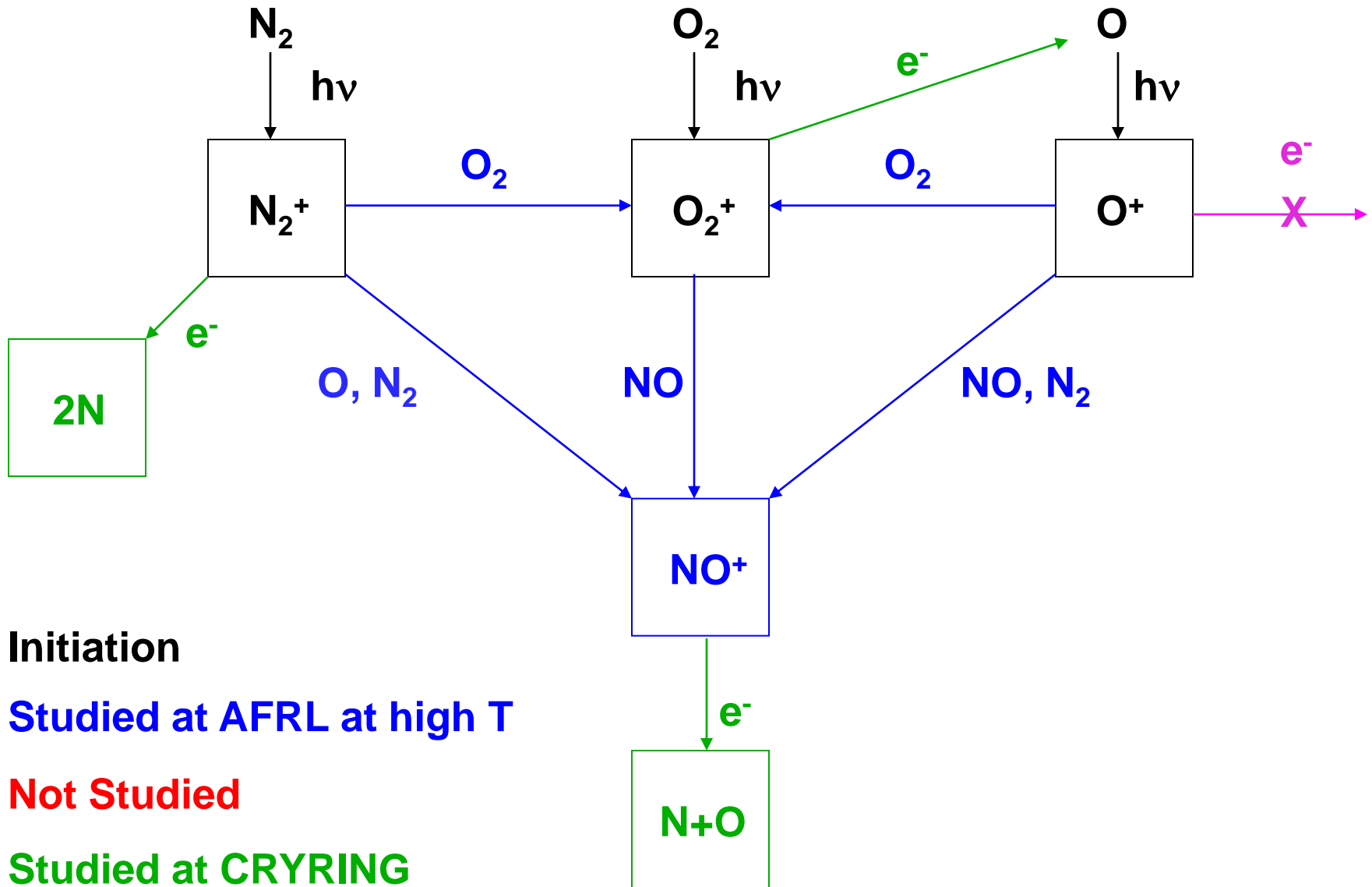
Outline



- **Brief introduction to ionospheric chemistry**
 - **Reasons for energetic oxygen studies**
- **Techniques for O, O₂(v), O₂(a ¹Δ_g), and O₃**
- **Data examples for each reactant**



Summary of Main Ionospheric Chemistry of N and O species



Initiation

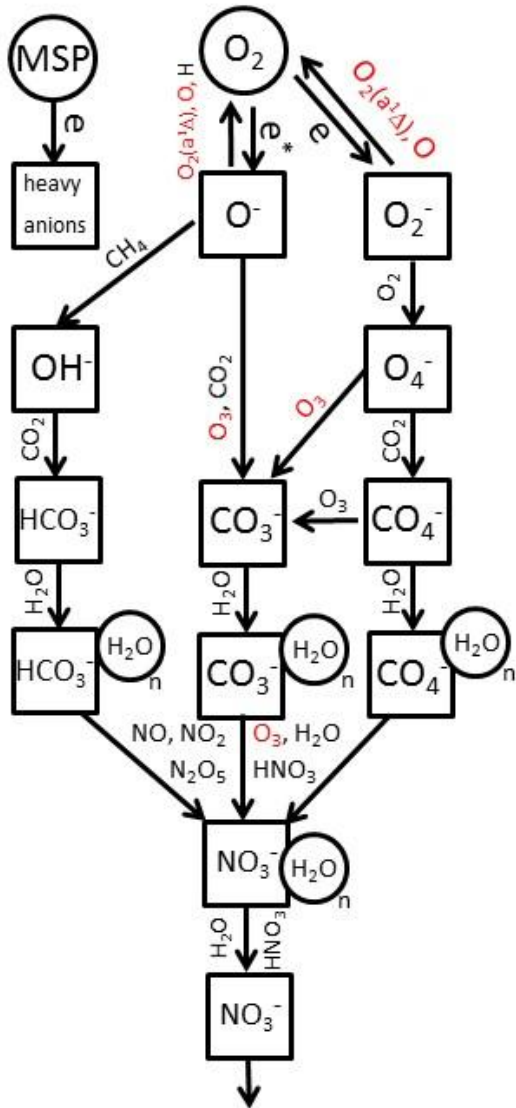
Studied at AFRL at high T

Not Studied

Studied at CRYRING



Negative Ion Reactions in the Atmosphere

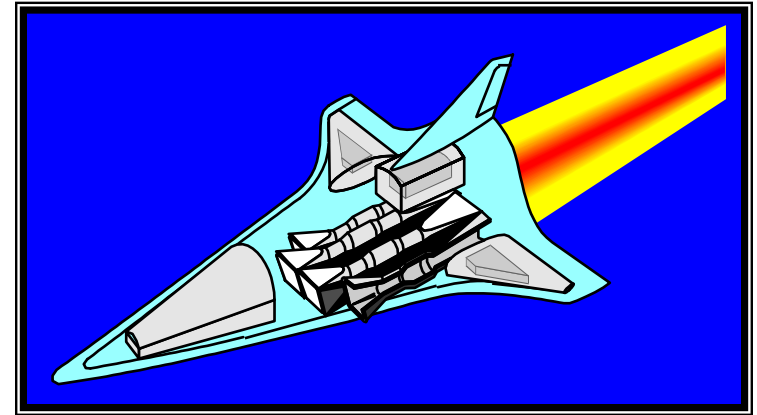
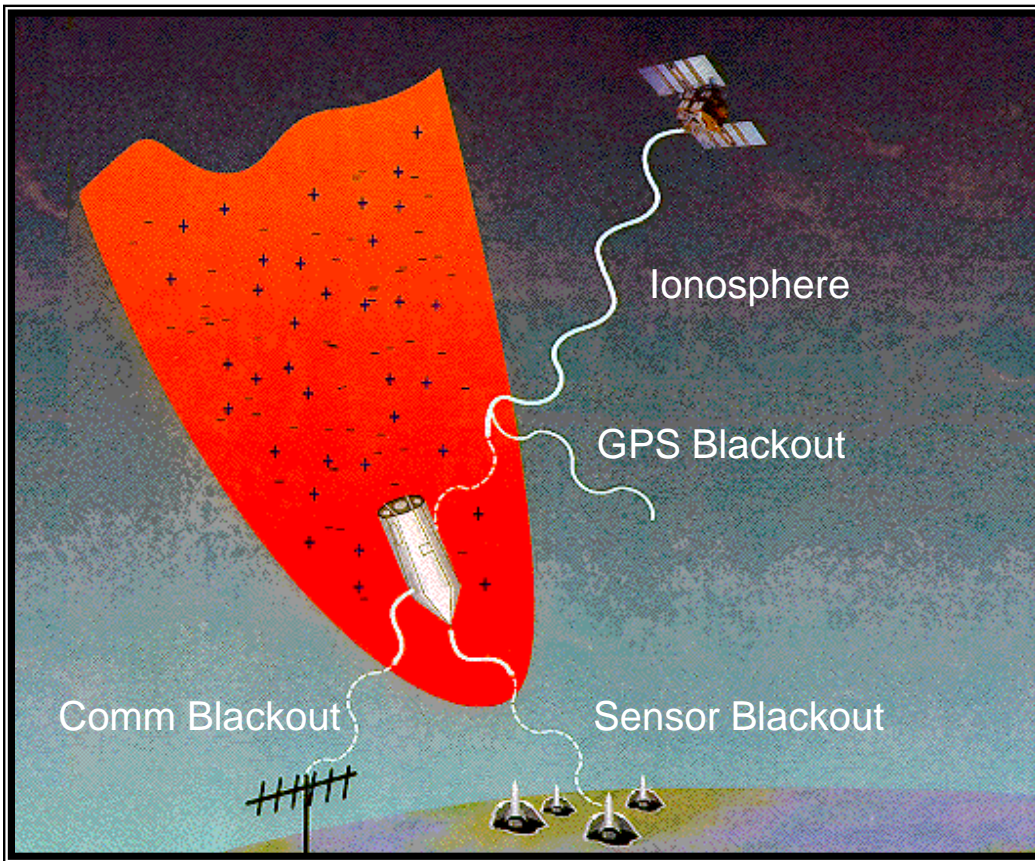


Numerous places reactive oxygen species are important

*For more information on atmospheric ion chemistry
Chem. Rev. 115, 4542–4570 (Feb 2015)*



Hypersonic Plasma Effects



- AJAX hypersonic concept vehicle utilizes air plasmas to aid combustion
- Plasma Blackout of C³I, GPS Navigation

**Combustor
Test @
Mach 2**

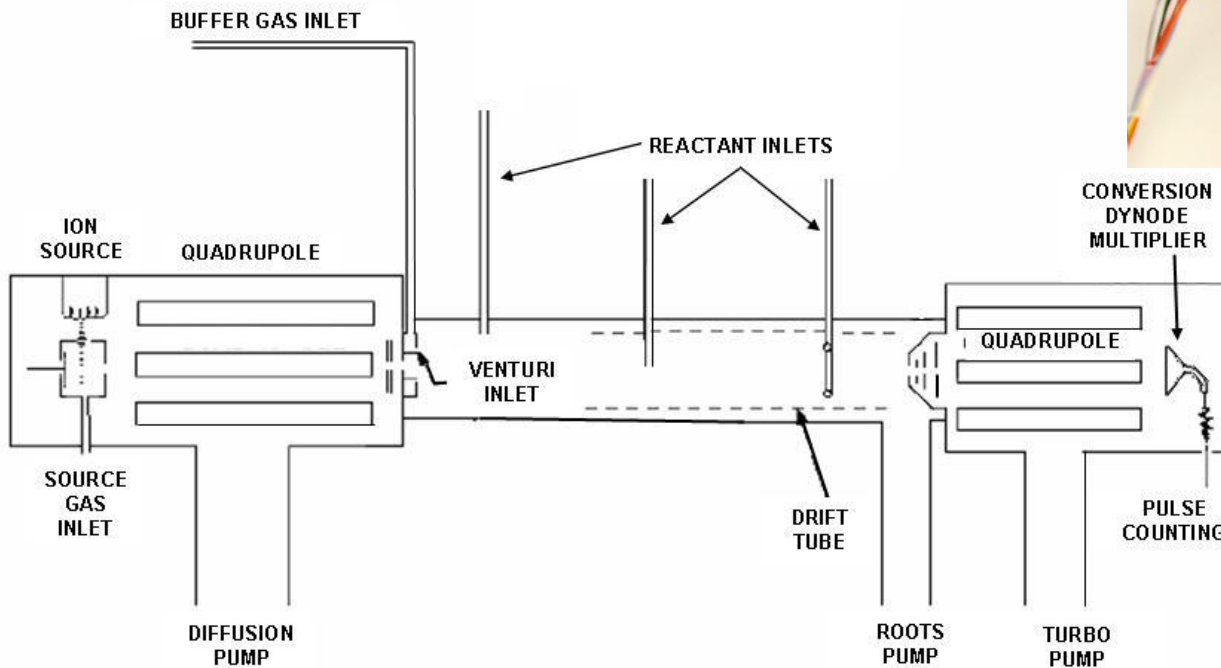
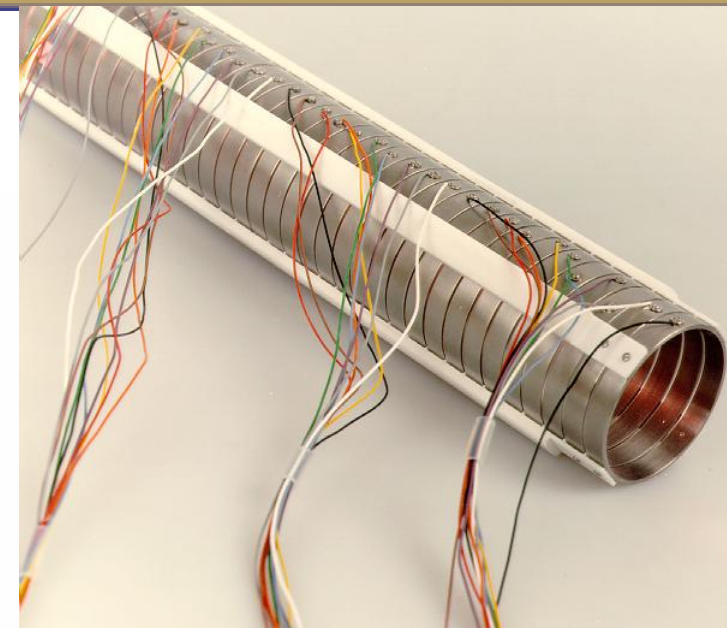




Selected Ion Flow Tube (SIFT) for $O_2(v)$



- T Range 85 - 550 K
- Pressure Range ~0.3 - 1 Torr
- Kinetic energy range 0.01 – 1eV

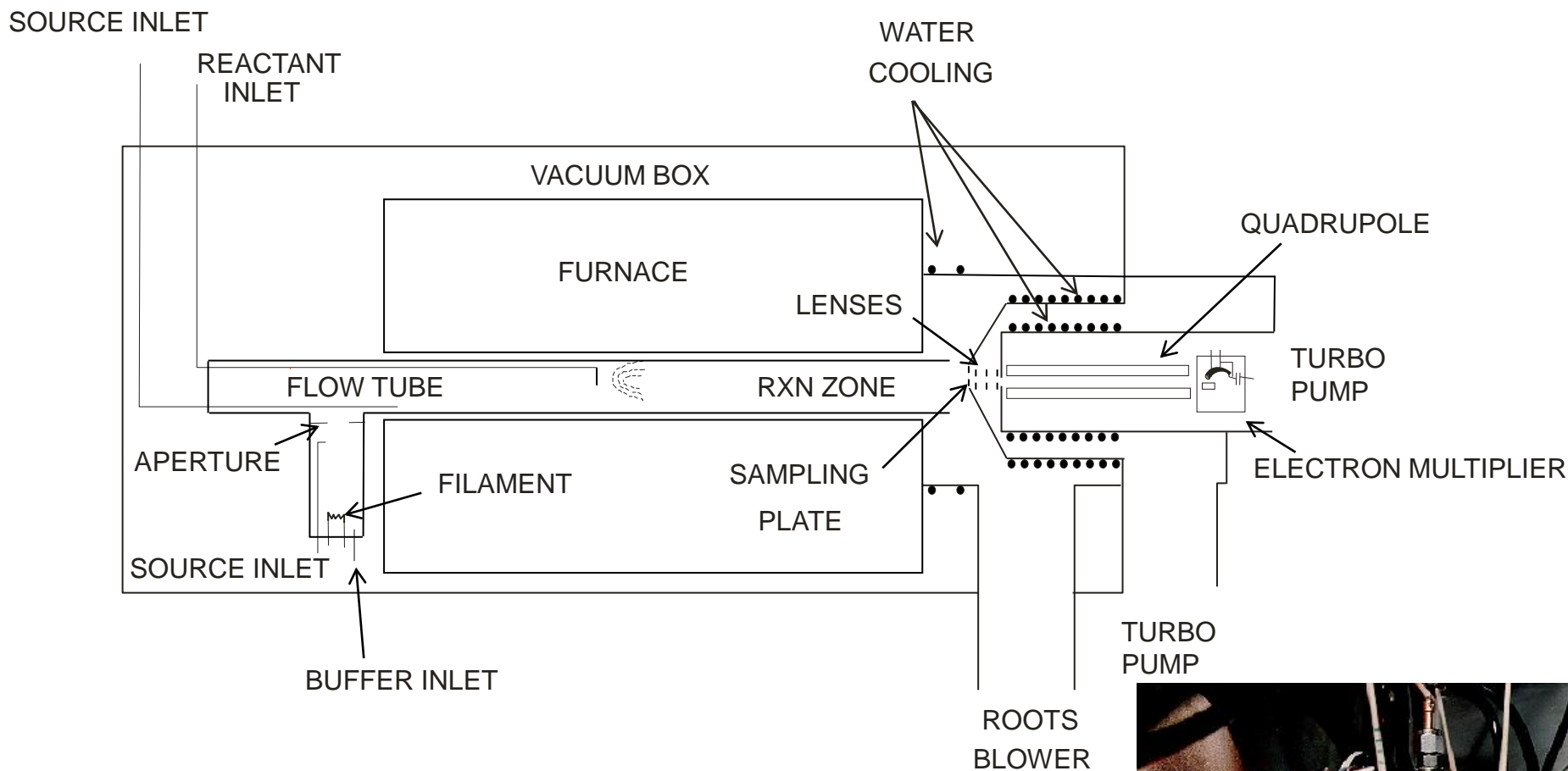


Important:

Translational energy distribution is Quasi-Boltzmann



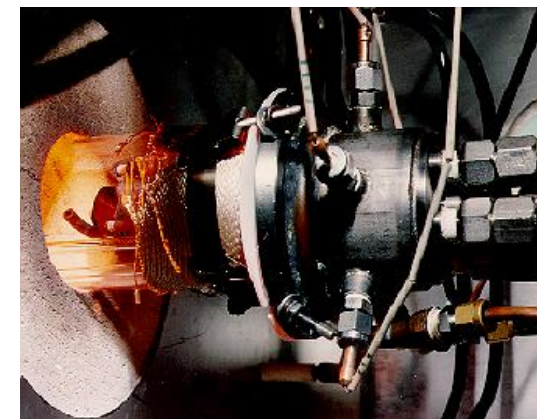
High Temperature Flowing Afterglow (HTFA)



Temperature Range 300-1800 K

Ceramic tube - 1800 K

Quartz tube - 1400 K



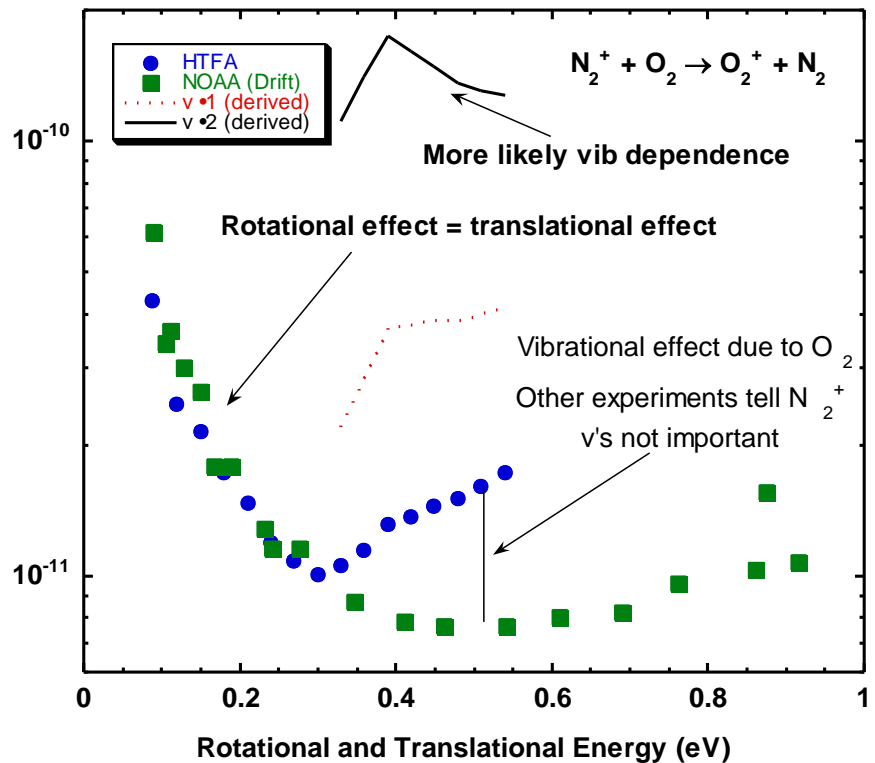
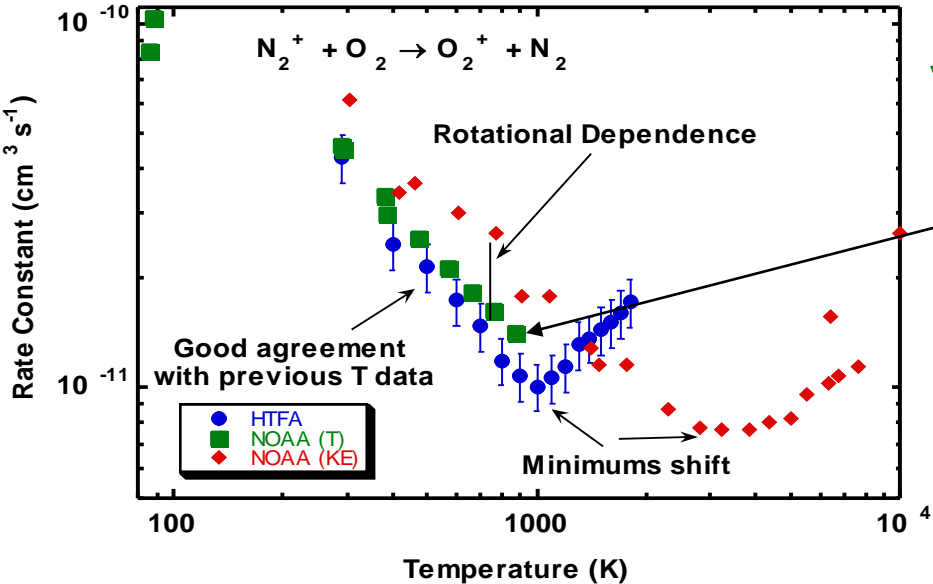


Data from HTFA and Drift Tube → Internal Energy Dependence



vs. T or KE

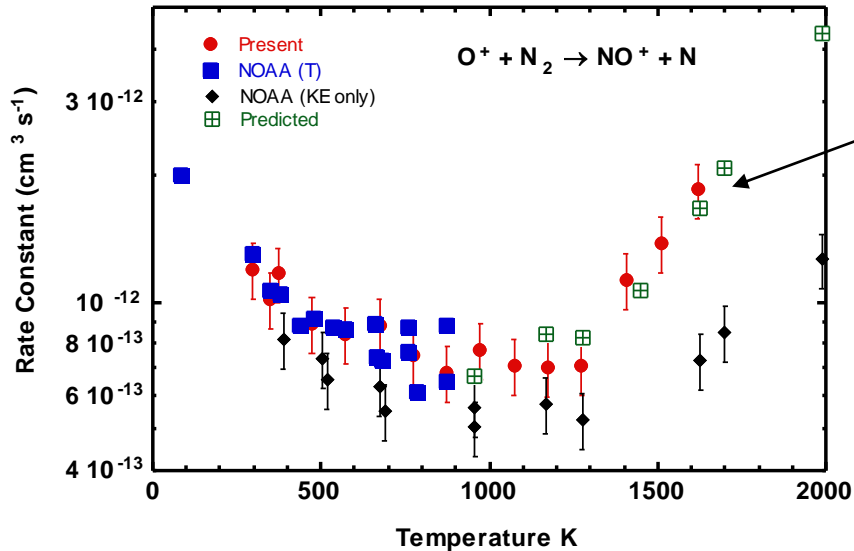
900 K limitation miss upturn



Add rotational energy

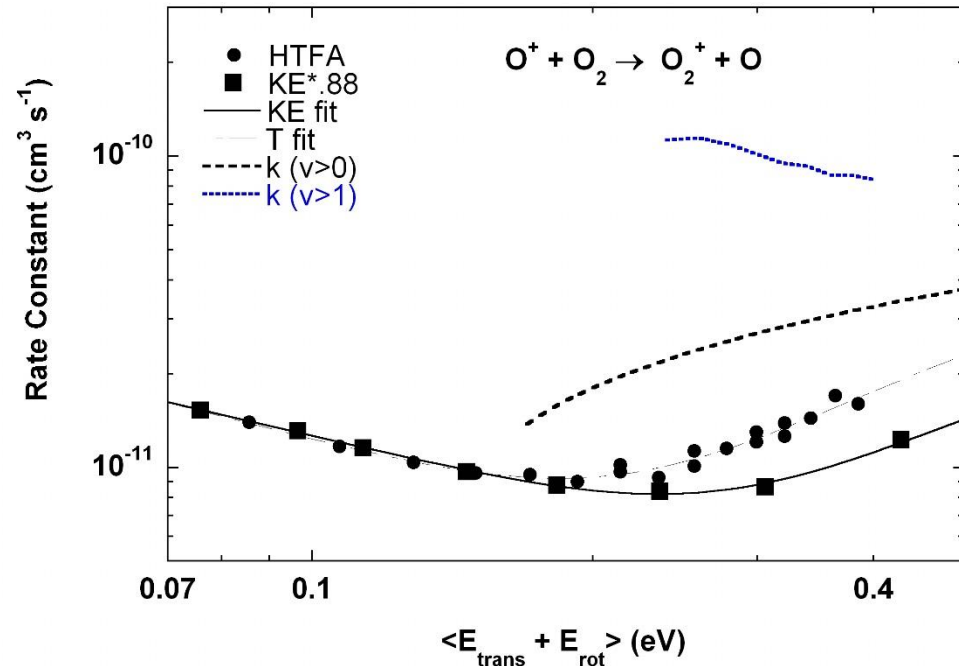


Most Important Ionospheric Reactions



Changes directions at 900 K
Upturn vs T mostly N_2 ($v \geq 2$)
Previous data missed upturns

Big difference between T and KE
Vibrationally excited O_2 important
maybe 10x faster for $v = 2$

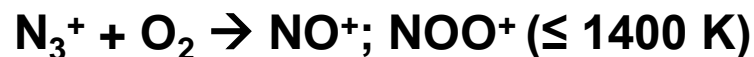
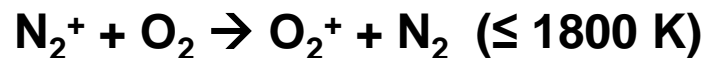
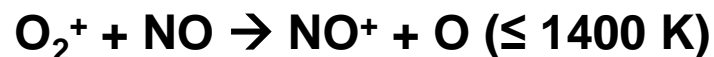
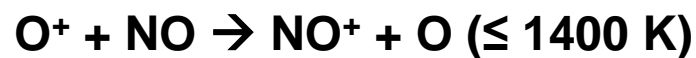
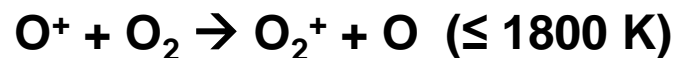




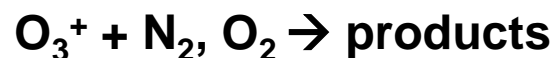
N, O Reactions Studied at AFRL by this technique



See: *Adv. in Gas Phase Ion Chem. vol. 4, p 85-136 (Dec 2001)*



Chemistry of NOO^+



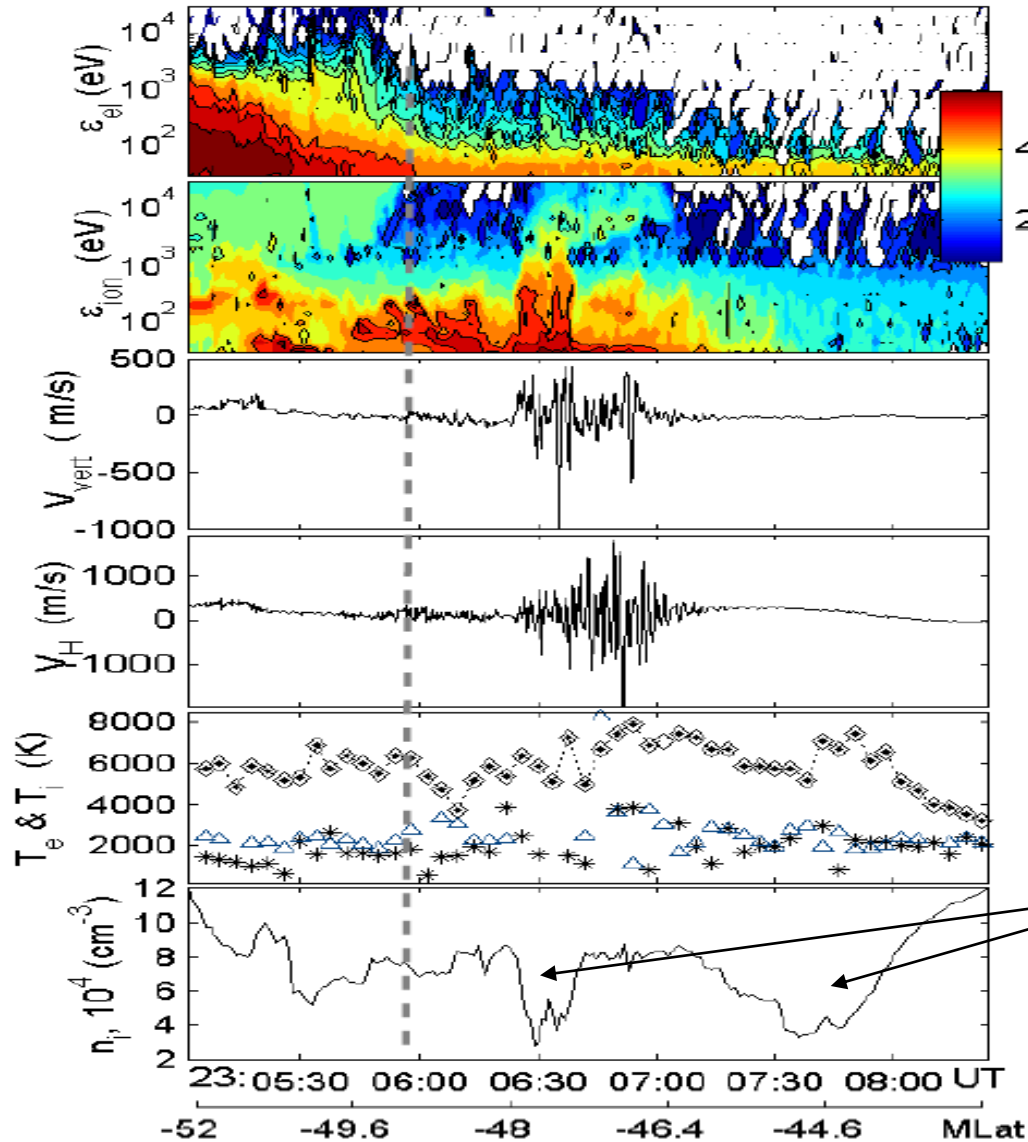


Ionospheric Data (with E. Mishin, B. Burke)



MISHIN ET AL.: STORMTIME SAPS-RELATED TROUGHS

SAPS wave structure
observed by DMSP F14 at
2307 UT on 6 April 2000



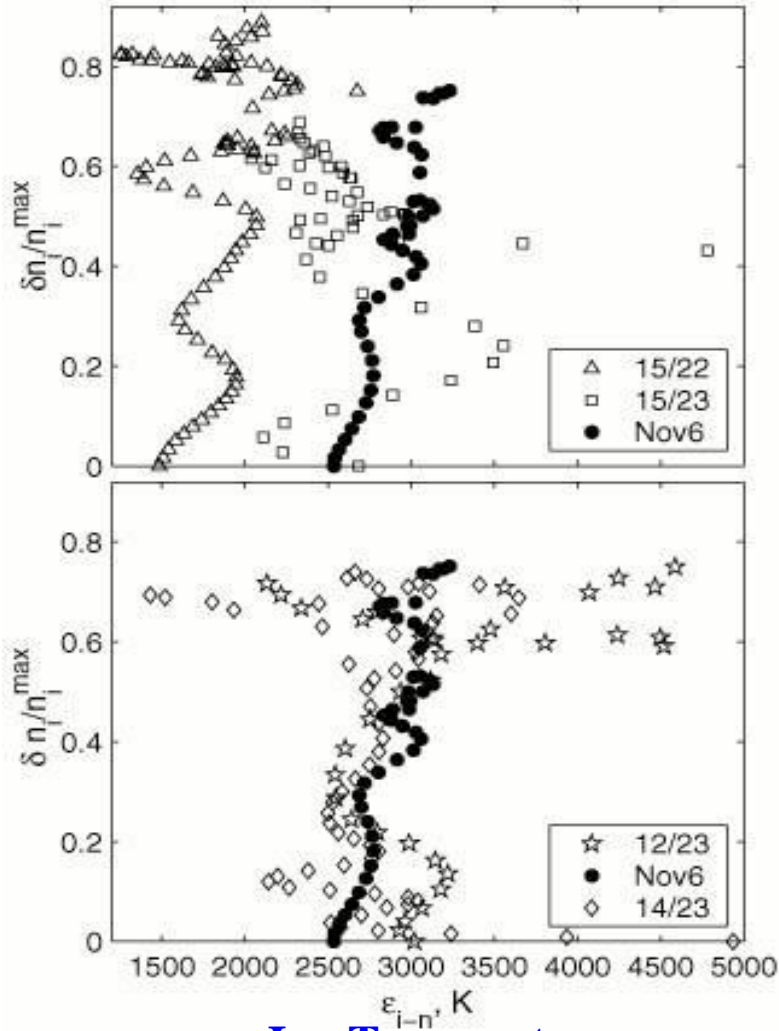
Density trough



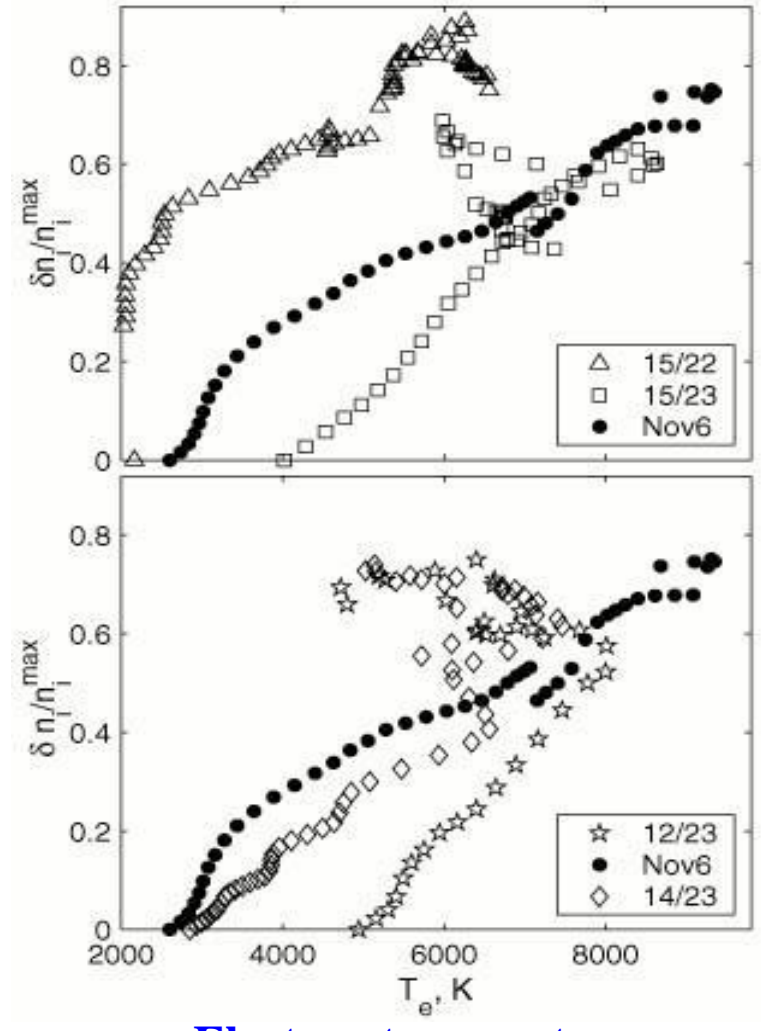
Ionospheric Depletion Data



$$(n_i - n_i^{\max}) / n_i^{\max}$$



Ion Temperature



Electron temperature

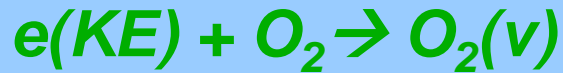
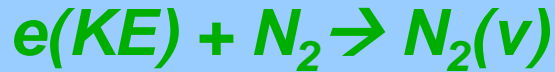
Dependence on electron temperature was not expected



Chemistry Leading to Trough



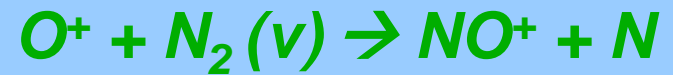
Vibrational excitation



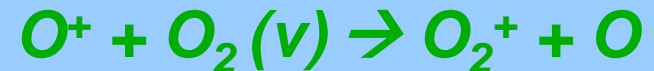
More KE \Rightarrow more vib.



Atom Transfer



Charge Exchange

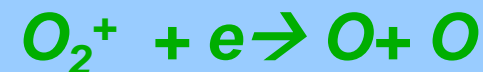


More vib. \Rightarrow k faster

Uses HTFA data



Dissociative Recombination



More diatomic \Rightarrow more recomb.

Determines the ionization balance in the F region

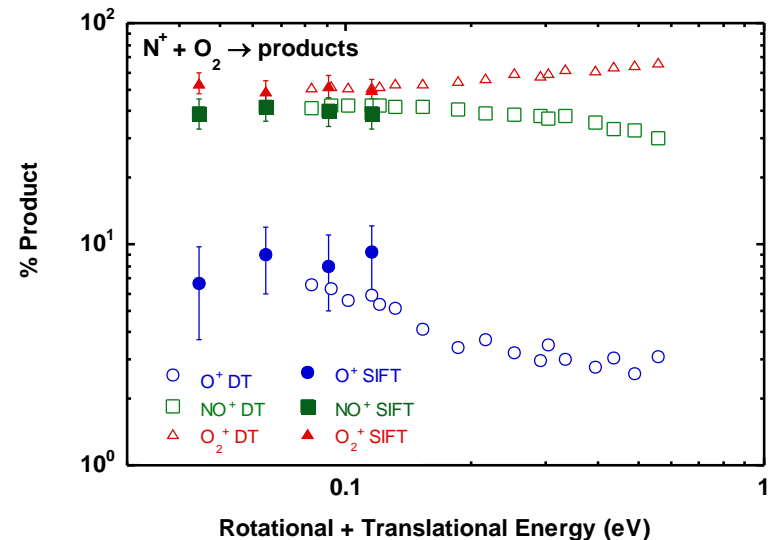
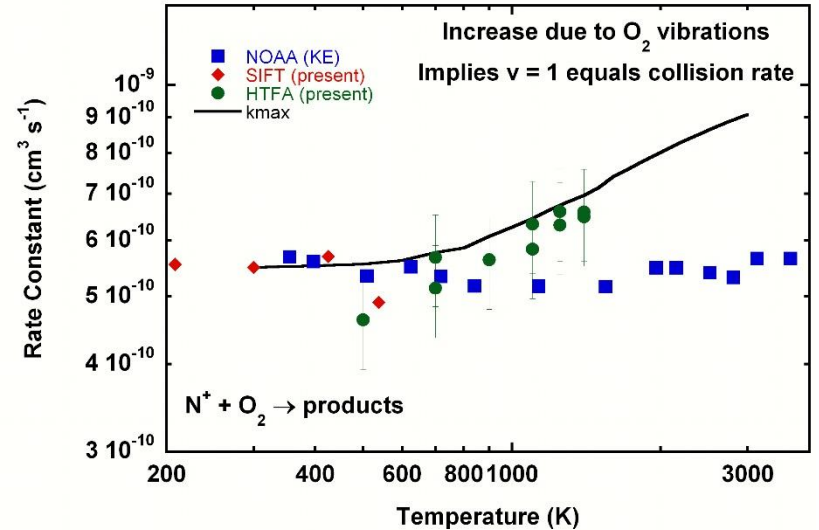
Our data shows that vibrational excitation leads to faster conversion of atomic species (non-recombining) to diatomic species (recombining)



N⁺ + O₂ Chemistry Studied

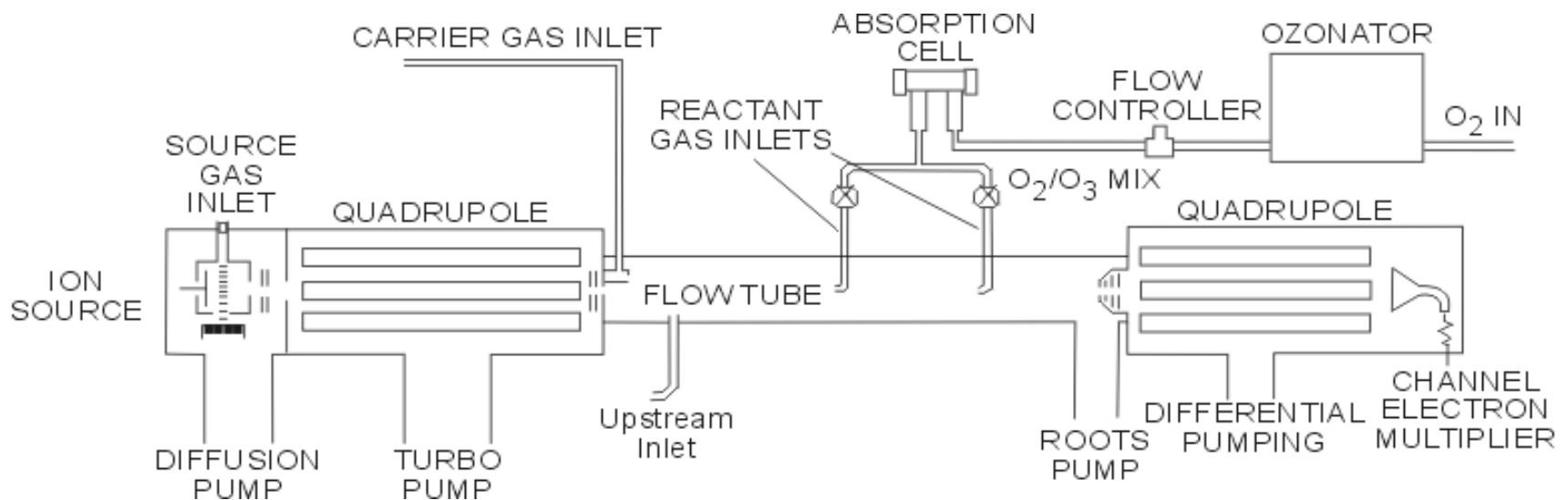
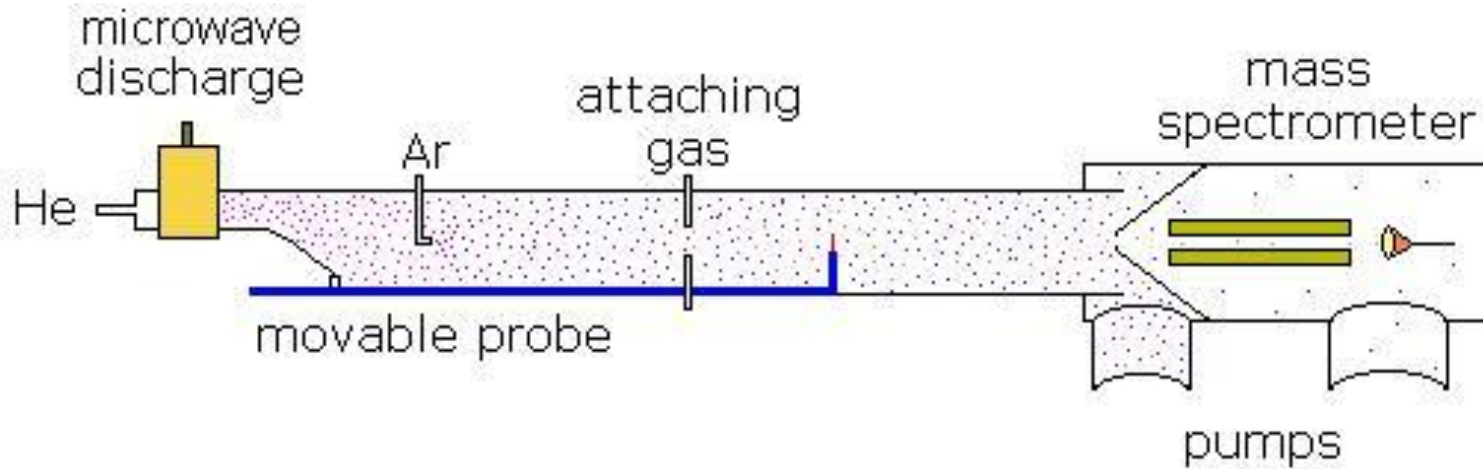


- Rate constants to 1400 K
 - Rotational and translational energy do little to change rate or products
 - $v = 1$ goes at k_{langevin}
- Three products formed
 - ~50% O₂⁺ (CT)
 - ~40% NO⁺
 - ~10% O⁺
- All NO⁺ is ground state (¹Σ) even though NO⁺ (³Σ) is exothermic (300- 550 K)





Instruments for ion and electron molecule studies of O_3





Atmospheric Ozone Reactions

J. Phys. Chem. A; 106(6), 997-1003 (Jan 2002)



TABLE 1: Reaction Rate Constants for Reactions of Ozone at 300 K Measured with the Selected Ion Flow Tube (SIFT)^a

reaction	products	$k; [k_c]$ ($10^{-9} \text{ cm}^3 \text{ s}^{-1}$)	branching fractions	$-\Delta H$ kJ/mol
$\text{O}^- + \text{O}_3 \rightarrow$		$1.7; [1.5]$		
	$\text{O}_3^- + \text{O}$		0.81	63
	$\text{O}_2^- + \text{O}_2$		0.19	294
$\text{O}_2^- + \text{O}_3 \rightarrow$	$\text{O}_3^- + \text{O}_2$	$1.3; [1.2]$	1.00	160
$\text{OH}^- + \text{O}_3 \rightarrow$		$1.4; [1.4]$		
	$\text{O}_3^- + \text{OH}$		0.90	28
	$\text{HO}_2^- + \text{O}_2$		0.08	100
	$\text{O}_2^- + \text{HO}_2$		0.02	47
$\text{NO}_2^- + \text{O}_3 \rightarrow$		$0.18; [1.1]$		
	$\text{NO}_3^- + \text{O}_2$		0.99	264
	$\text{O}_3^- + \text{NO}_2$		0.01	-14
$\text{NO}_3^- + \text{O}_3 \rightarrow$	no reaction	$<0.005; [0.97]$		
	$\text{NO}_2^- + 2\text{O}_2$			21
$\text{CO}_3^- + \text{O}_3 \rightarrow$	no reaction	$<0.001; [0.98]$		
	$\text{O}_2^- + \text{CO}_2 + \text{O}_2$			98
$\text{CO}_4^- + \text{O}_3 \rightarrow$		$0.46; [0.93]$		
	$\text{O}_3^- + \text{CO}_2 + \text{O}_2$		0.93	81
	$\text{CO}_3^- + 2\text{O}_2$		0.07	108

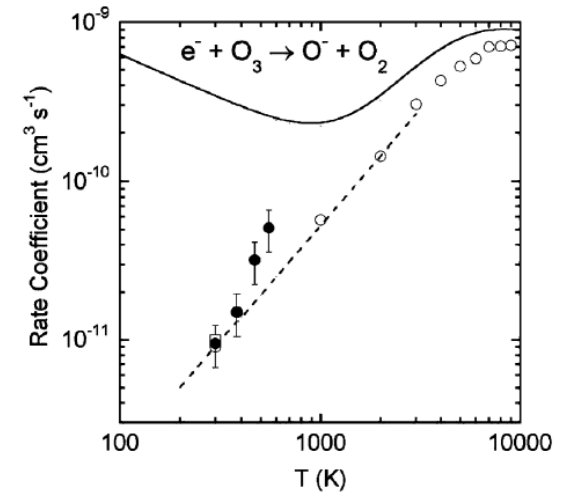
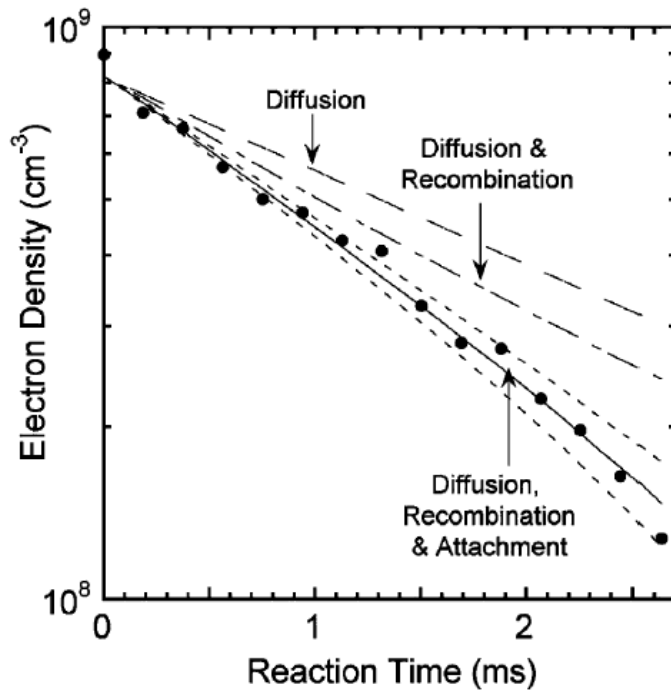
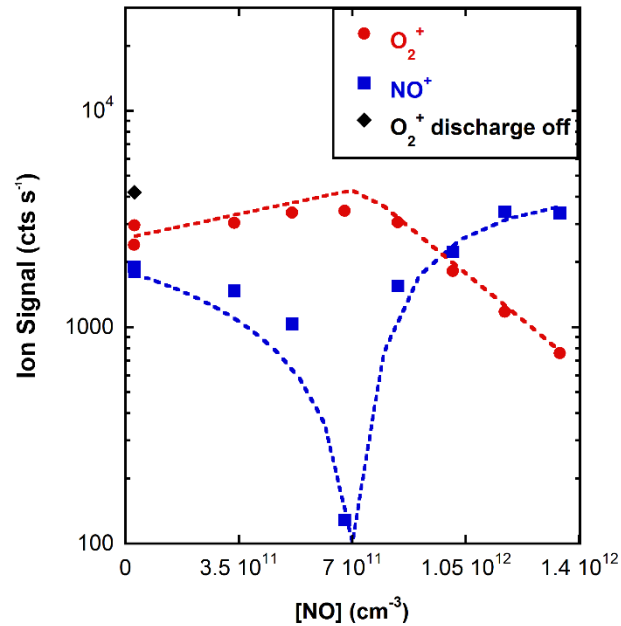
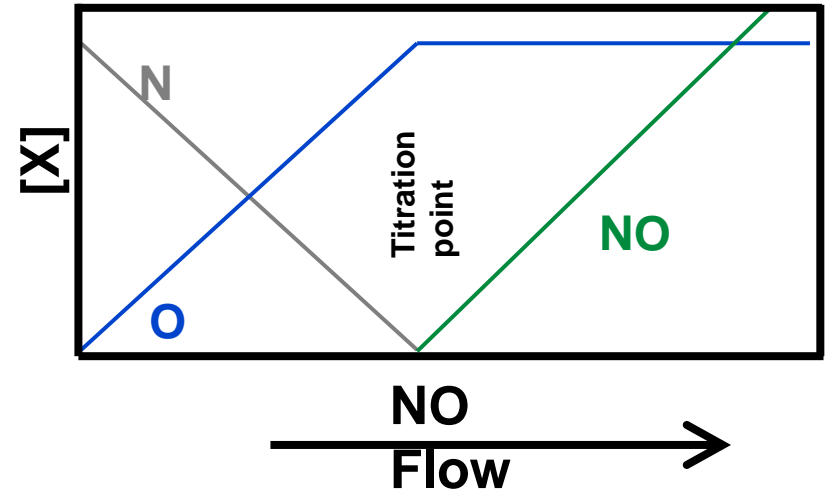
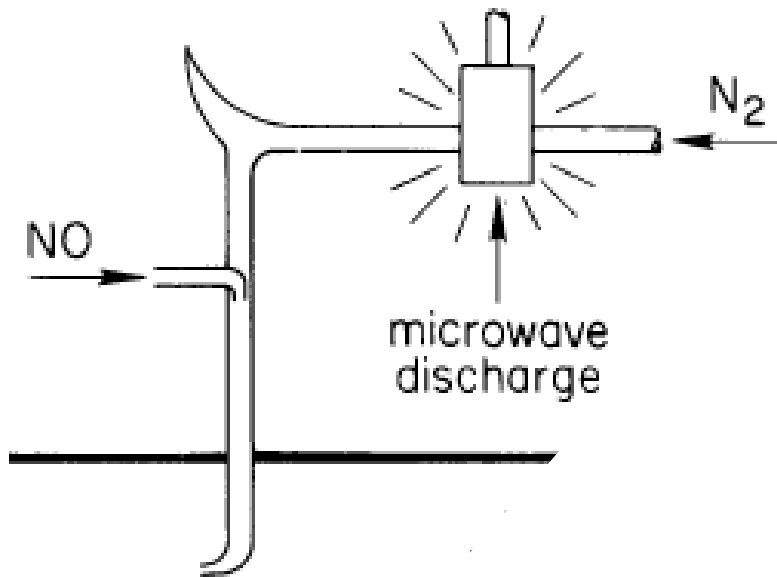


FIG. 2. Electron attachment rate coefficient versus temperature. Present results (●) and swarm upper limit of Fehsenfeld *et al.* [3] (□) are true thermal values. The remaining data are plotted versus electron temperature. The drift tube results of Stelman *et al.* [4] (dashed line) are derived from a least squares fit of the combined data for 200 and 300 K rovibrational temperature ozone reacting with energetic electrons. The electron beam results of Skalny *et al.* [8] (○) and Senn *et al.* [2] (solid line) were derived by those authors from measurements of the reaction cross section for 300 K rovibrational temperature ozone with energetic electrons.

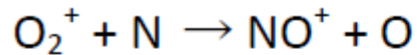


O, N atom reaction technique

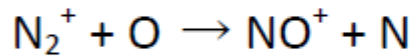




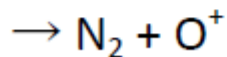
Positive Ion - Atom Results



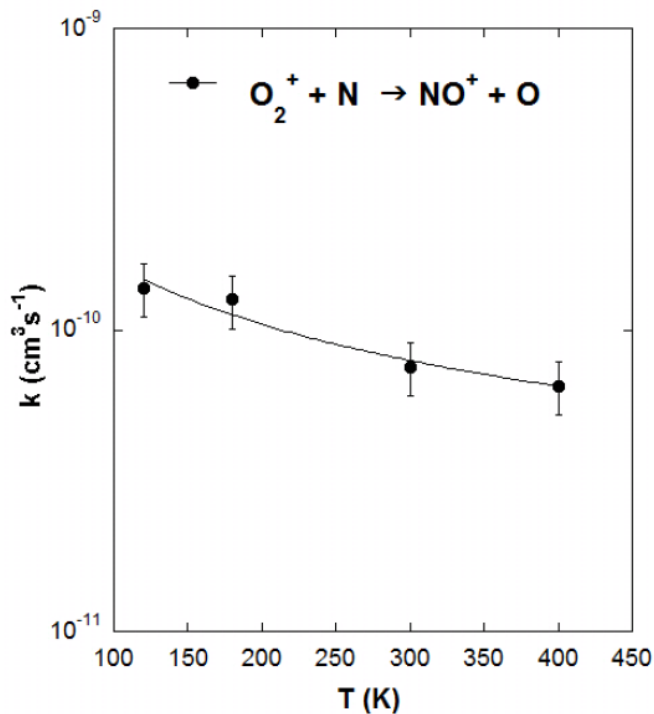
$$\Delta H_r^0(0K) = -96.5 \text{ kcal/mol (1)}$$



$$\Delta H_r^0(0K) = -70.6 \text{ kcal/mol (2a)}$$



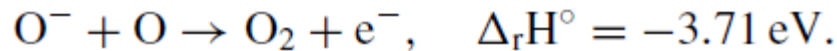
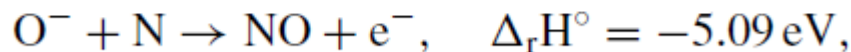
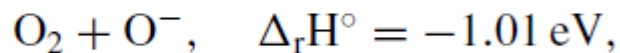
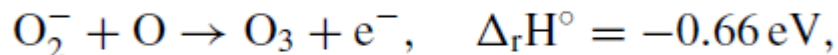
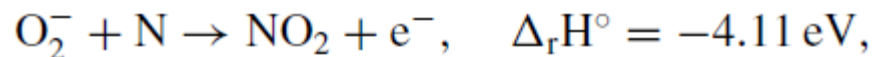
$$\Delta H_r^0(0K) = -45.2 \text{ kcal/mol (2b),}$$



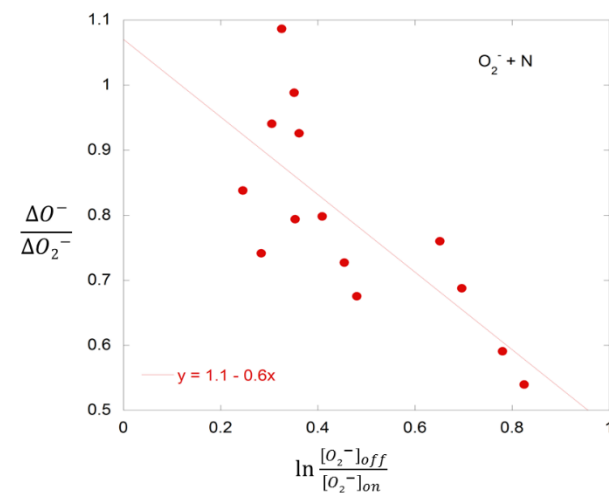
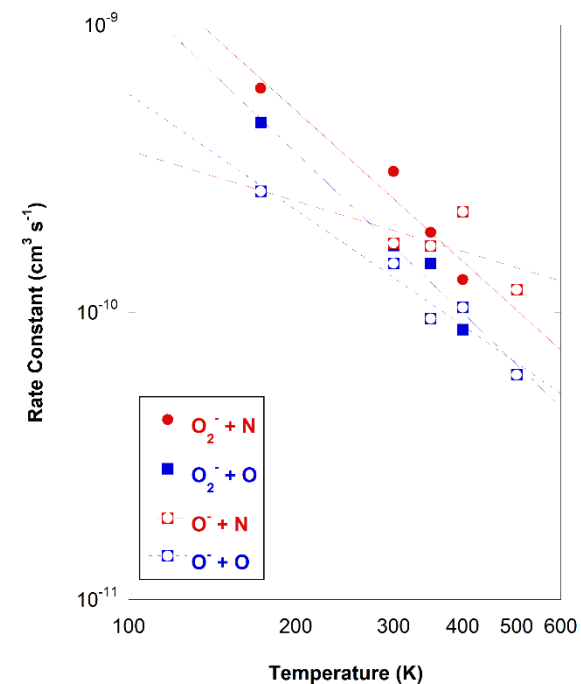
J. Chem. Phys. **142** DOI: 154305
(Apr 2015)



Negative Ion - Atom Results

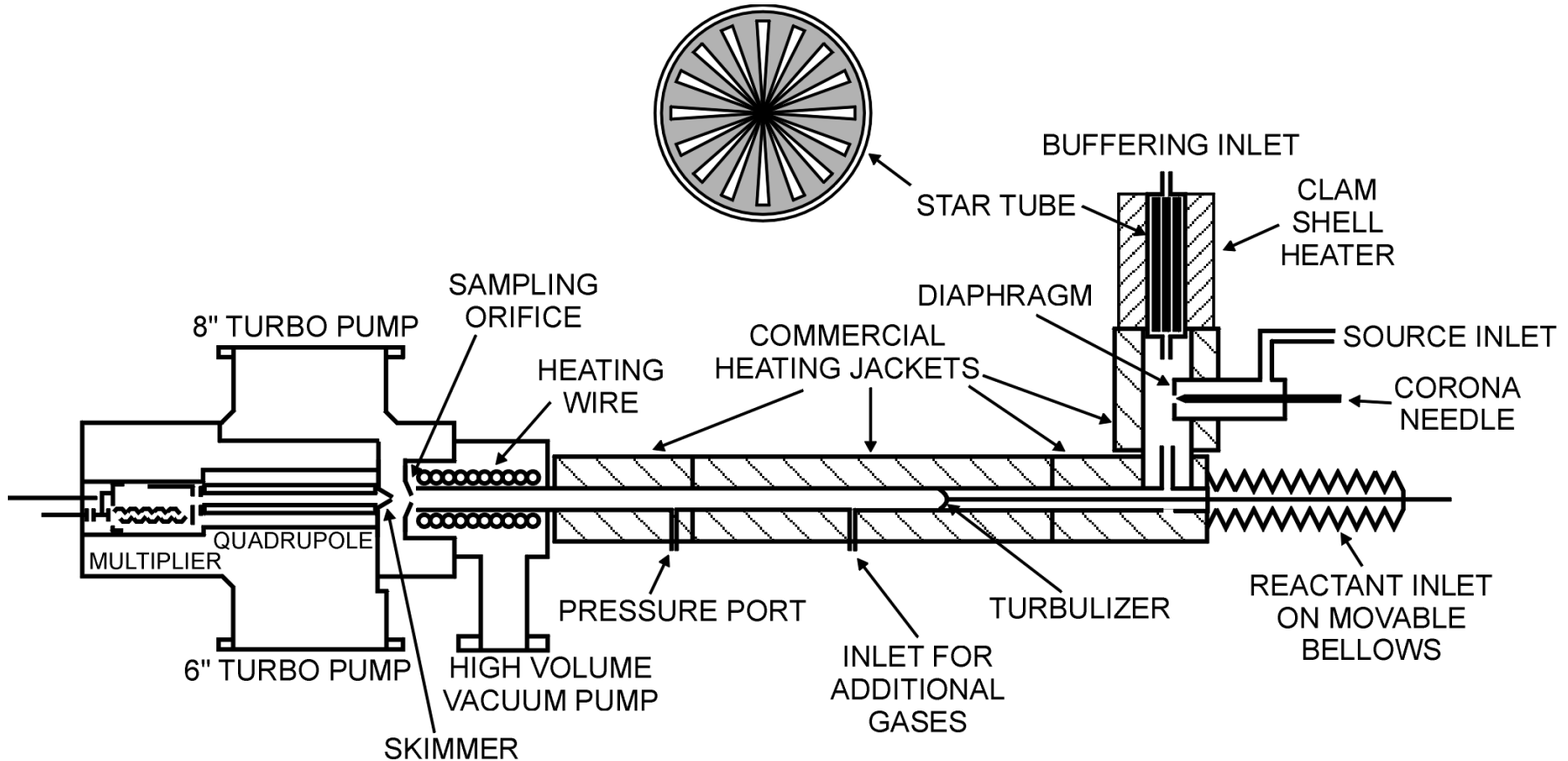


J. Chem. Phys. **139**, 144302,
doi: 10.1063/1.4824018 (Oct 2013)





Turbulent Ion Flow Tube



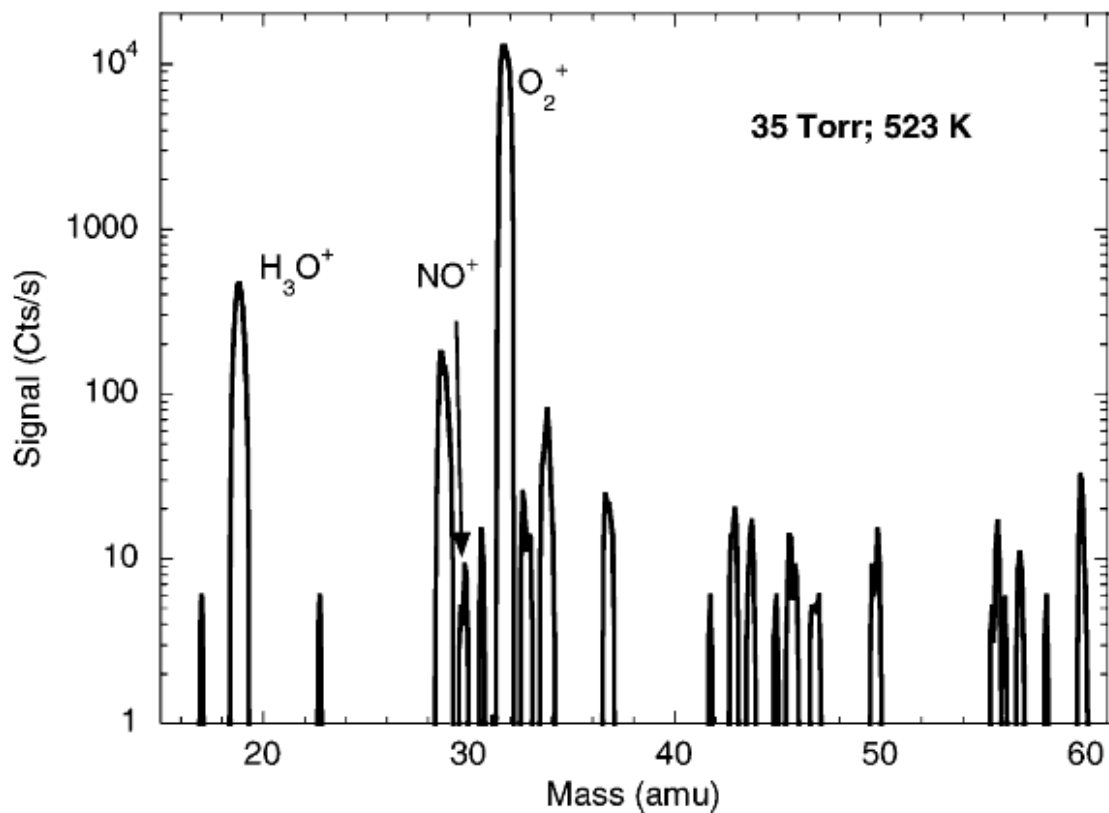
10-760 Torr

300-700 K

Also for very slow rate constants



Limits for O_2^+ with N_2



T (K)	k (upper limit)
423	2×10^{-21}
523	4×10^{-21}
623	1×10^{-20}

Figure 1. Mass spectrum taken at 35 Torr and 523 K.



$O_2(a^1\Delta_g)$ background



- $O_2(a^1\Delta_g)$ emissions at 1270 nm contribute to the IR airglow¹
- $O_2(a^1\Delta_g)$ created in O_2 discharges²
 - Electron impact on O_2
 - $O_2(b^1\Sigma_g^+)$ collisional quenching by O_2
- Affects oxidation chemistry
 - Materials processing³
 - Oxygen-iodine lasers⁴

¹Handbook of Geophys. & Space Environ., A.S. Jursa, ed. 1985. ³Jeong et al., *Plasma Sources Sci. Technol.* **7**, 282-285 (1998)

²Sweitzer and Schmidt, *Chem. Rev.* **103**, 1685-1757 (2003)

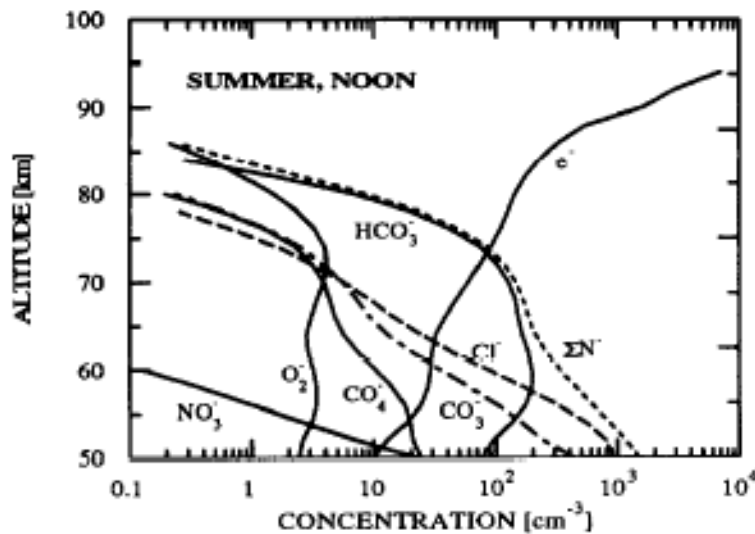
⁴A. P. Naparatovich et al., *J. Phys. D: Appl. Phys.*, **34**, 1827 (2001)



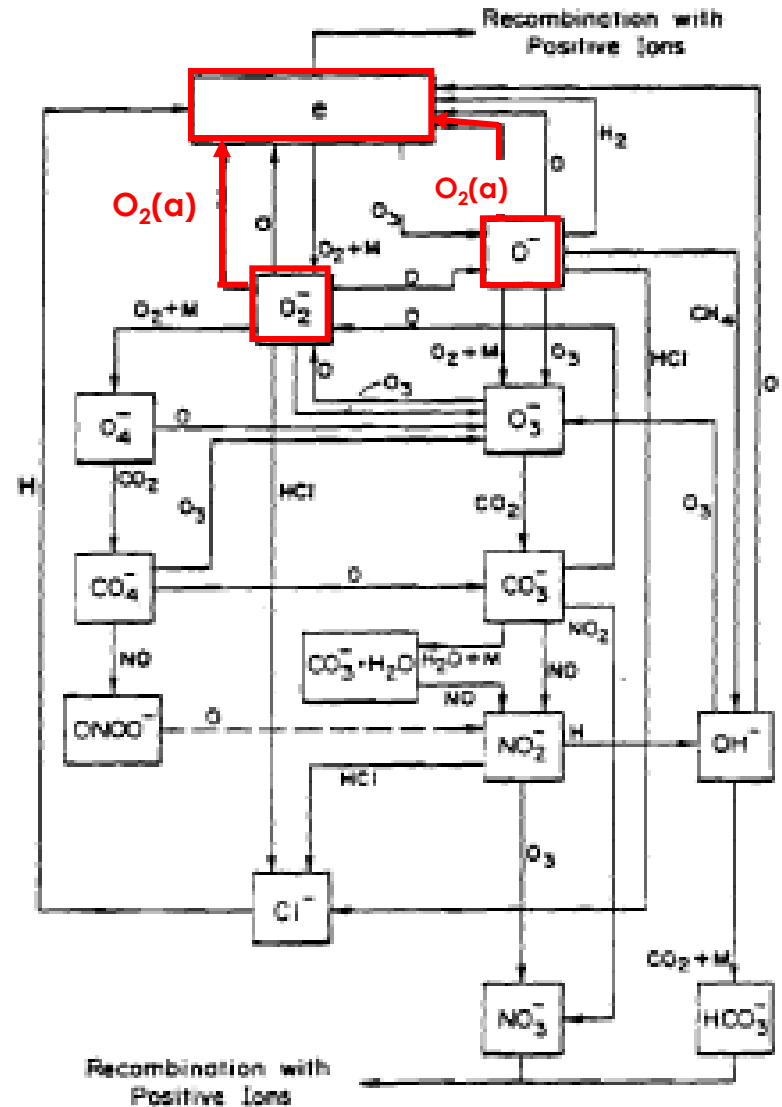
$O_2(a^1\Delta_g)$ in Ionosphere D-region



- Reactions controlling e- concentrations in ionosphere influence radiowave propagation¹



Brasseur and DeBats, *JGR*, **91**, 4025 (1985)



¹Handbook of Geophys. & Space Environ., A.S. Jursa, ed. 198



Previous Kinetics for $O_2(a^1\Delta_g)$ Reactions



- Large disparity in the literature values of the rate constants for O^- and $O_2^- + O_2(a^1\Delta_g)$
- $O_2^- + O_2(a^1\Delta_g)$
 - NOAA flowing afterglow (FA)¹⁰ $2.0 \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$
 - Upschulte et al., FA¹¹ $2.4 \times 10^{-11} \text{ cm}^3 \text{ s}^{-1}$
- $O^- + O_2(a^1\Delta_g)$:
 - NOAA, FA $3.0 \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$
 - Upschulte et al., FA $3.3 \times 10^{-11} \text{ cm}^3 \text{ s}^{-1}$
 - Belostostky et al., plasma model $1.9 \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$
 - Stoffels et al., plasma model $1.3 \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$

¹¹JPC. 98, 837 (1994)



AFRL Experiments



- Utilize newly designed $O_2(a^1\Delta_g)$ emission detection scheme to re-measure the kinetics for the O^- , $O_2^- + O_2(a^1\Delta_g)$ reactions from 200-700 K
- Calibrate detection setup vs. absolute standard
 - Settle the discrepancy in the literature values
- Expand the studies to other ion-molecule reactions with $O_2(a^1\Delta_g)$

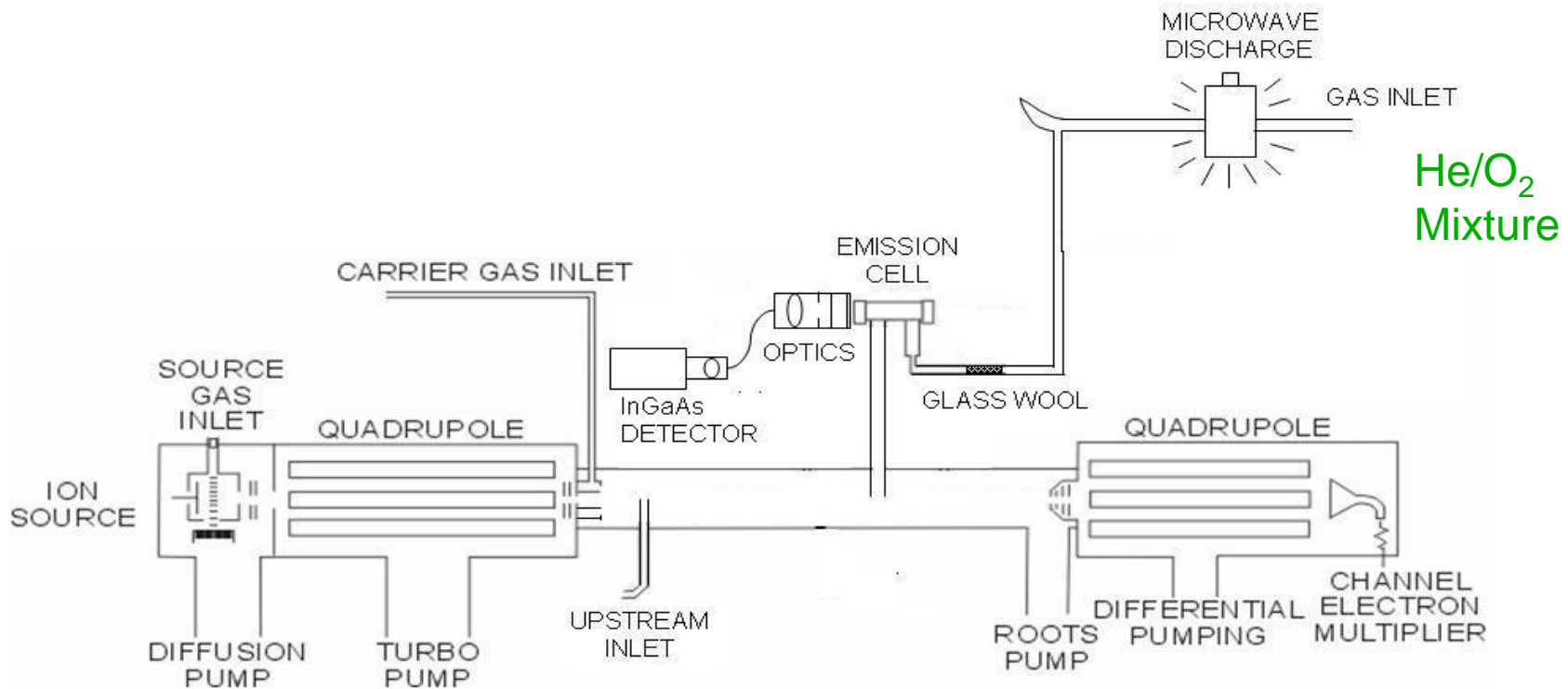
***J. Phys. Chem. A.* 111, 5218-5222 (June 2007)**

***J. Phys. Chem A* 112, 3040-3045 (Apr 2008)**



SIFT with $O_2(a^1\Delta_g)$ Detection

Initial experiments

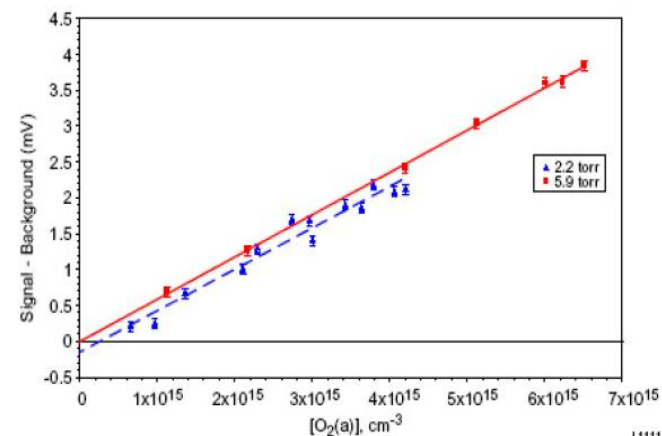
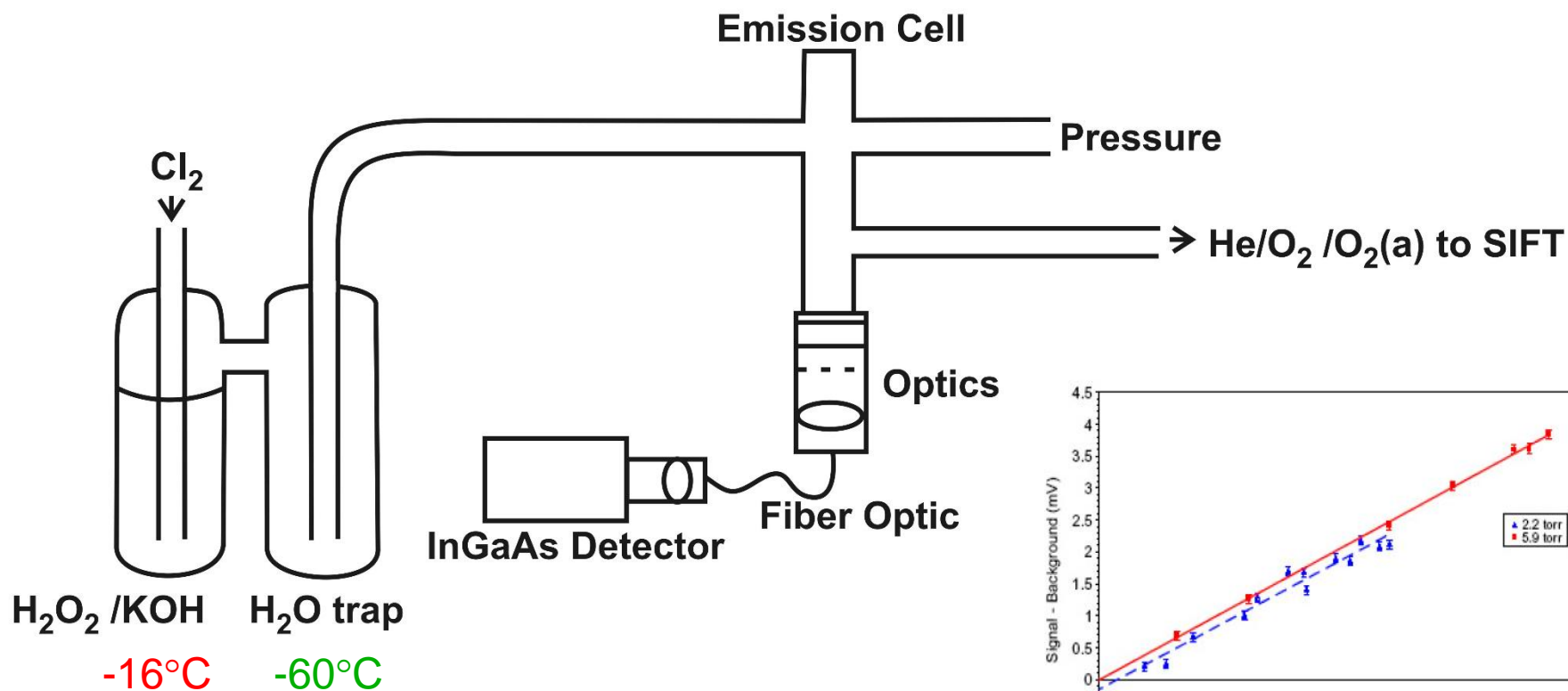


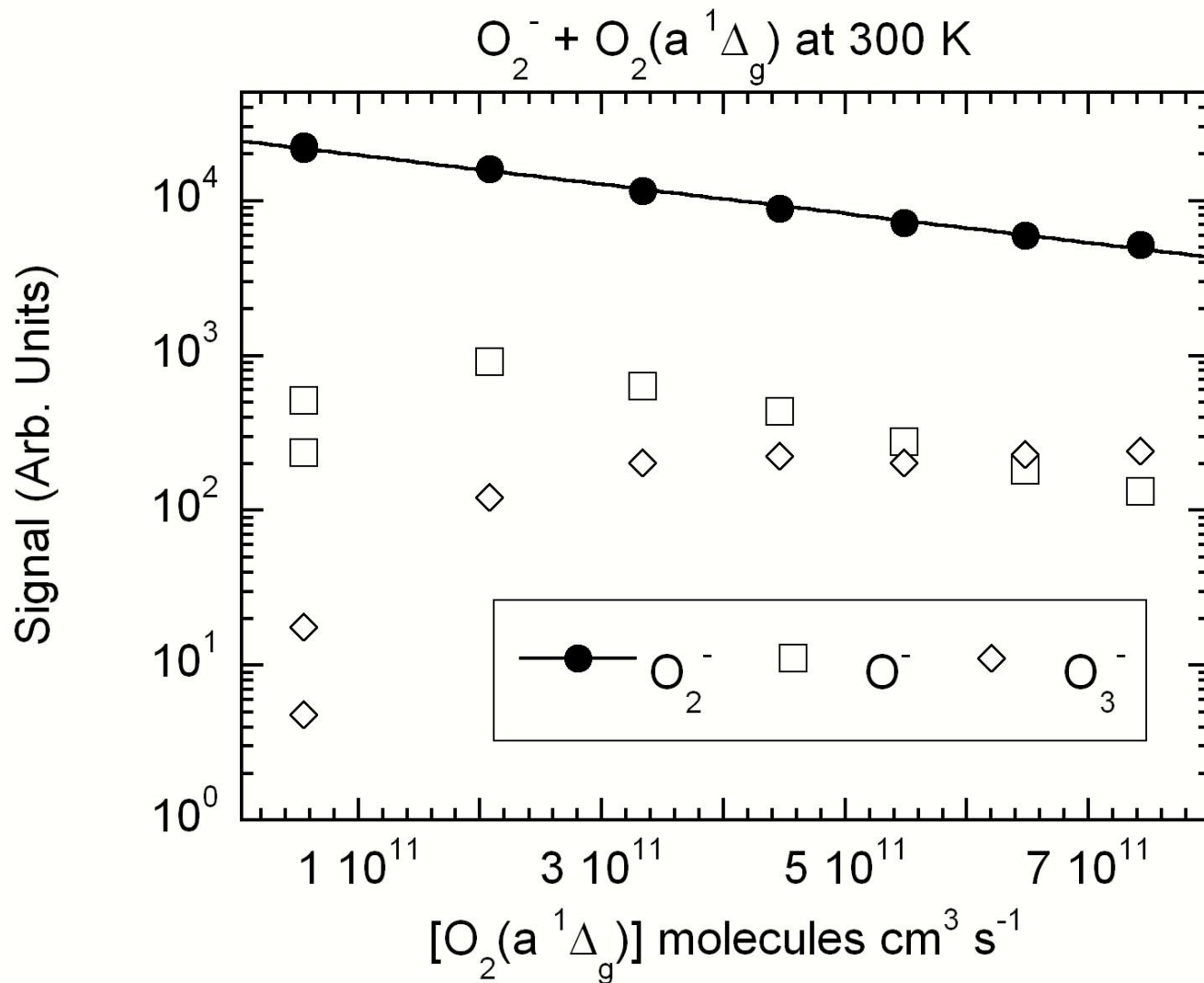
Glass wool: quenches most O atoms from discharge

Typical % of total O_2 concentration in SIFT: 9% $O_2(a^1\Delta_g)$, 1% O, <1% O_3



Chemical O₂(a ¹Δ_g) Generation Can Study Temperature Dependencies





Need to account for O, O₃ impurities



$O_2^- + O_2(a^1\Delta_g)$ 298 K Results



- $k = 6.6 \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$
 - 90% of collision rate constant \therefore very efficient
 - **3x higher than previous highest values**
- Upshulte et al. FA experiments
 - Kinetics data showed fast and slow decay
 - Incorrectly assumed slow decay was correct
 - New bi-exponential fit shows fast decay = NOAA value
 - O_2 source gas present in FA flow tube
 - **Electrons present re-attach to O_2 \therefore lower apparent decay in FA measurements**



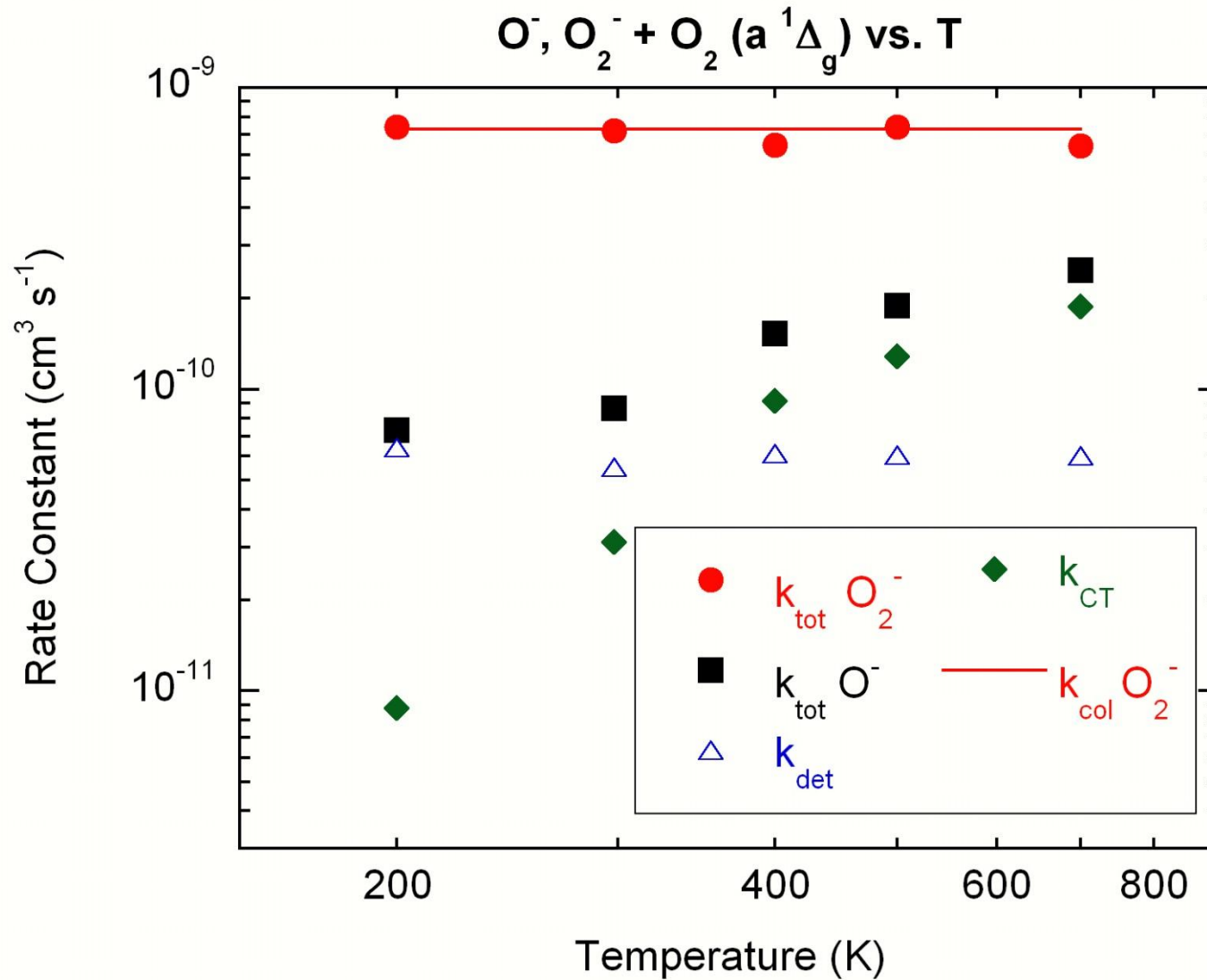
$O^- + O_2(a^1\Delta_g)$ 298 K Results



- $k = 1.1 \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$
 - 12% of $k_{\text{col}} = 9.0 \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$
 - **2-3x < previous values (even more – see T work)**
- **New charge transfer product channel observed:**
$$O^- + O_2(a^1\Delta_g) \rightarrow O_3 + e + 60 \text{ kJ mol}^{-1} \quad <70\%$$
$$\rightarrow O_2^- + O + e - 3 \text{ kJ mol}^{-1} \quad >30\%$$
- **New channel important pathway in low-pressure O_2 discharges where:**
 - O^- is primary ion
 - Three-body O_2^- formation negligible



Temperature Dependencies 200-700 K





Implications for Models Temperature Dependencies



- **O^- rate constant at 298 K = $8.6 \times 10^{-11} \text{ cm}^3 \text{ s}^{-1}$**
- **D-region is cold $\therefore O^-$ rate constants are lower than previously assumed at low temperatures given positive T dependence**
- **Increased fraction of O^- converted to O_2^- at high temperatures increases conversion rate to electrons**
 - **Additional O_2^- rapidly converted to e^-**



Conclusions



Innovative techniques allow wide range of species studied

New techniques often yield unexpected results



Acknowledgments



In-house

Tom Miller

Skip Williams

Bob Morris

Sue Arnold

Nick Shuman

Shaun Ard

Oscar Martinez

Justin Wiens

Jenny Sanchez

Theory

E. Mishin - atmospheric modeling

Jurgen Troe – statistical modeling

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